Solar System overview



2) spin/orbit/shape

3) heated by the Sun

4) how do we find out



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Inventory (cont'd)

Many moons & rings

Mercury: 0 Venus: 0 Earth: 1 (1700km) Mars: 2 (~10km) Jupiter: 63 + rings Saturn: 60 + rings Uranus: 27 + rings Neptune: 13 + rings

Even among dwarf planets, asteroids, Kuiper belt objects, and comets. E.g.,

Pluto: 3 Eris: 1

Moons of Mars: Deimos & Phobos, ~10km



Moons of Jupiter

4 Galilean satellites (Ganymede, Callisto, Io & Europa),

~10³ km (close to Jupiter, likely primordial)





2001J3: 4km



Inventory (cont'd)

- ~10⁵ known small objects in the
- Asteroid belt (Ceres ~300 km)
- Kuiper belt (Eris, Pluto, Sedna, Quaoar, ~1000 km)
- Estimated: $\sim 10^{12}$ comets in the
- Oort cloud (~ 10⁴ AU)
- Associated:
- zodiacal dust

(fire-works on the sky: comets & meteorites)



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What are planets?

IAU (for solar system):

Orbits Sun, massive enough to be round and to have cleared its neighbourhood.

More general:

1) no nuclear fusion (not even deuterium): $T_c < 10^6 \text{ K}$ 2) pressure provided by electron degeneracy and/or Coulomb force (I ~ h/p ~ d) (d ~ atomic radius)



Orbits

inclination: largely coplanar (history) direction: all the same eccentricity: a few percent (except for Mercury)



Titus-Bode (fitting) law (1766)

planetary orbits appear to (almost) satisfy a single relation 'Predict' the existence of the asteroid belt (1801: Ceres discovered) coincidence or something deeper?

other systems?



Spin (obliquity)

smaller planets: almost random, affected by impacts and giant planets

Real giant planets (J&S): ~aligned with orbit, stable



Shape --- the bigger the rounder

All gaseous planets are spherical. Large rocky objects are rather spherical. Smaller ones are less so.

The Moon (~1700km)	an asteroid (~50km)				
	in the second		h	R	g=GM/R ²
and the second se			(km)	(km)	(m/s²)
		Earth	8	6400	9.8
and a second		Mars	24	3400	3.7

scaling: highest mountain on Earth ~8 km (on Mars ~ 24 km) **h** * **g** ~ constant rough estimate: irregular body has mountain h ~ R ==> R ~ 240 km thus: objects with R > 240 km are approximately spherical



Passively Heated by the Sun --- the further the cooler

Typically we observe objects in reflected light, however, all objects emit re-processed thermal radiation which is observable at longer wavelengths.

Blackbody temperature for a non-self-luminous spherical body at distance *a* away from the Sun (with albedo *A* -- reflectivity)

$$L_{abs} = (1-A) \frac{\pi R_{\rho}^{2}}{4\pi a^{2}} 4\pi R_{s}^{2} \sigma T_{s}^{4}; \quad L_{em} = 4\pi R_{\rho}^{2} \sigma T_{\rho}^{4}$$
If $L_{abs} = L_{em}$, then $T_{\rho} = \left| \frac{R_{o}}{2a} \right|^{1/2} T_{s} (1-A)^{1/4}$

$$a (AU) \qquad A \qquad T_{pred}(K) \qquad T_{act}(K)$$
Mercury 0.4 0.06 422 K 100-725 (?)
Venus 0.7 0.77 230K 733 (?)
Earth 1 0.30 255K 288 (?)
Mars 1.5 0.25 218K 223 good
Jupiter 5 0.51 113K 125 (?)
Saturn 9 0.47 83K 95 (?)
Uranus 19 0.51 60K 60 good
Neptune 30 0.62 40K 60 (?)

How do we know?

presence: orbit: angular size; occultation of a star; radar signal strength; size: lander; blackbody+albedo orbits of moons; perturbation on other planets; artificial moon mass: rotation: magnetic field: core: surface composition: rings: INTERSTELLAR WIND . . . HELIOPAUSE Notable planetary missions HELIOSPHERE Voyager 1 & 2 are now formally **VOYAGER 1** Interstellar Missions interstellar material outside heliopause

Cassini is currently visiting Saturn & dropped Huygens probe on Titan sea on the surface of Titan?







The Cassini-Huygens Mission





Phobos (@Mars): 7.6 hr Moon (@Earth): 1 month Charon (@Pluto): 6 day

What about retro-grade moons?